



12th IPWG Workshop

Microphysical Controls of High-Frequency Microwave Scattering

Implications for Radiance Simulations, O-B Biases, and MP Bias Correction

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ROADMAP

From Microphysical Sensitivity to MP Bias Correction

All-sky microwave DA hinges on unbiased O-B; conventional VarBC corrects only air-mass and scan bias, yet in cloudy scenes microphysical assumptions inject the TB variability.

1 Sensitivity to MP

Brief Reviews on Paper Points

Kim et al. (2025), IEEE TGRS

“Sensitivity of Microwave RT Simulations to Cloud Microphysics Assumptions”

Frozen density & habit drive TB sensitivity at scattering channels.



2 Implication for O-B

THIS WORK

Shin & Kim, IPWG-12 — this study

The predictors driving MP-injected TB variability: XGBoost + SHAP rank



3 Proposing MP bias correction

THIS WORK

Shin & Kim, IPWG-12 — this study (proposed)

Extend airmass / scan bias correction with MP-aware predictors.

Microphysics Changes Microwave TB

General RTM forward-simulation pipeline — from model state to top-of-atmosphere brightness temperature



- MP enters the RT through two pathways:
 - via NWP input fields
 - via the RTM's PSD, particle density & habit
- **Coupling problem:** both routes change together, so we isolate each with controlled RTM runs.
- **What we test:** TB sensitivity to individual MP factors — PSD, particle density & habit — at fixed water content.

Isolating each microphysical lever



EXP#1

Rainwater PSD



EXP#2

Frozen-particle PSD



EXP#3

Frozen density



EXP#4 (REF)

Full MP assumptions,
but fixed fields

PSD — the Self-Limiting Lever

RTM pathway — particle-size distribution. Both rain & frozen PSD lose their grip at high frequency.



EXP#1 · RAINWATER PSD

6 K → 0.6 K

self-limiting (10 → 166 GHz)

- Large at 10 GHz — large-particle emission dominates rainwater extinction
- Collapses with frequency as small/large extinction and emission/scatter offset each other
- N_0 & μ matter more than the distribution family (EXPD $\approx$ DMD)



EXP#2 · FROZEN-PARTICLE PSD

2.6 – 5.3 K

comparable, then constrained

- Super-exponential Field PSDs push mass to small particles → weaker depression
- Forward-scatter cancels: high g (→ 0.78) at 166 GHz offsets extreme depression
- Net effect: variability shrinks at high frequency, like rainwater PSD

Bottom line: PSD shape is a second-order lever — pick a reasonable family and move on. The real levers are density & the input field.

Figures = TB sensitivity — the max-min spread of TB depression (ΔTB) across the swapped MP assumptions, shown by GMI channel (10–183 GHz).

Density & Field — What Matters

The dominant RTM lever (density) and the coupling NWP lever (input field) — where TB accuracy is won or lost.



EXP#3 · FROZEN-PARTICLE DENSITY

9.3 K · 49 K extreme

the dominant lever

- Controls Q_{sca} far more than any PSD choice (SPHERE / COX88 / MG08)
- Sets how strongly frozen hydrometeors scatter at high frequency
- **Get this right first — the single most important RTM assumption**



EXP#4 · INPUT FIELD × MODULE

up to 16.4 K

couples the whole chain

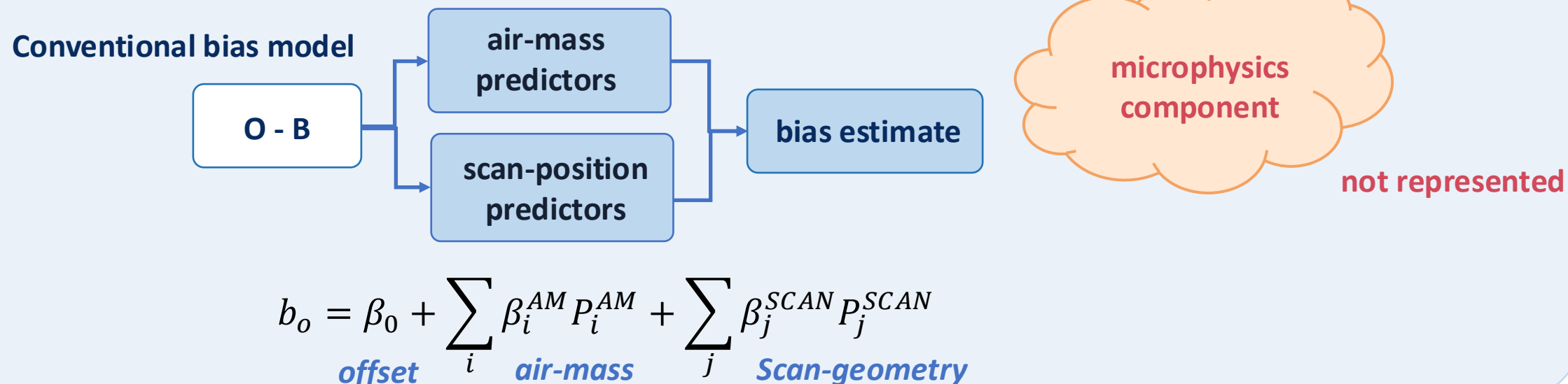
- Case A (WDM6 rain): RTM's DMD $\mu=1$ suppresses dTB 9.3 → 7 K at 10 GHz
- **Case B (THOM frozen): MP + Input fields resonates → 166 GHz spread 7.7 → 16.4 K (>2×)**
- **NWP field & RTM module are not independent — consistency wins**

Bottom line: density dominates the module and the field can more than double the spread — accuracy comes from a physically consistent chain, not one 'realistic' component.

Figures = TB sensitivity — the max-min spread of TB depression (ΔTB) across the swapped MP assumptions, shown by GMI channel (10–183 GHz).

Implication for O-B biases

Motivation: all-sky O-B departures contains more than air mass and scan bias



Conventional VarBC predictors were designed mainly for slowly varying systematic errors.

air-mass predictors:

$$P^{AM} = \{TPW, T_s, \Gamma, T_{850}, T_{200}\}$$

scan-position predictors:

$$P^{SCAN} = \{\theta, \text{Scan Position}, FOV\}$$

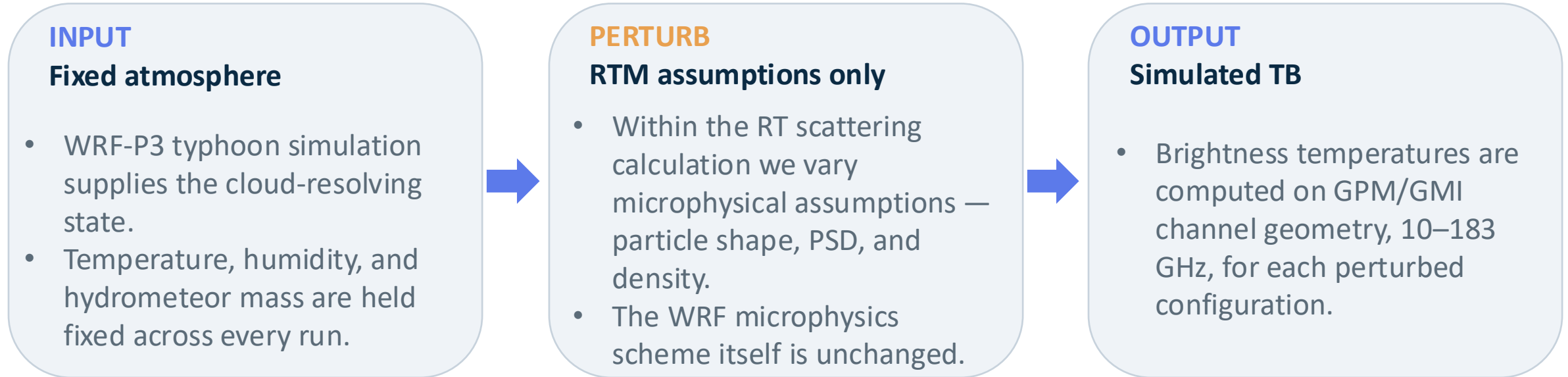
Key implication

- If MP assumptions alter $H(x)$, their effect can appear directly in O-B departures.
- Microphysics-induced TB variability — rain & ice PSD, habit, density — has no predictor of its own.
- It is aliased into the air-mass & scan coefficients, *quietly contaminating the correction.*

Finding the predictors driving MP-induced TB variability

Fixed-State RTM Experiments

Perturbing microphysical assumptions inside the radiative transfer — the atmospheric state is held fixed



“By holding the atmospheric input fields fixed, the experiment isolates the component of TB variability attributable to microphysical assumptions within the radiative-transfer simulation.”

→ TB differences reflect RTM microphysics assumptions — not differences in dynamics or water content.

Microphysics-Perturbation Sets

1 — 2 — 3

Each set is a controlled perturbation of one MP factor — isolating which predictor family the bias model must include



Shape / habit

Default · ARTS-OPT · SSRGA-OPT · SSRGA-RIME Poster 9.15



Rain PSD

RAIN_WSM6 · RAIN_WDM6



Ice PSD

ICE_PSDWSM6 · ICE_PSDTHOM · ICE_PSDMORR



Ice density

ICE_DENWSM6 · ICE_DENTHOM



CLW PSD

CLW_WSM6 · CLW_THOM · CLW_MORR



Two reference aggregates

MIE — all non-shape factors (shape variability excluded)

ALL — full MP-driven TB variability

P9.15 (TR) Integration of optimal habit selection with DDA and SSRGA models for improved microwave radiative transfer simulations (Kim and Shin)

WRF-P3 typhoon simulation cases

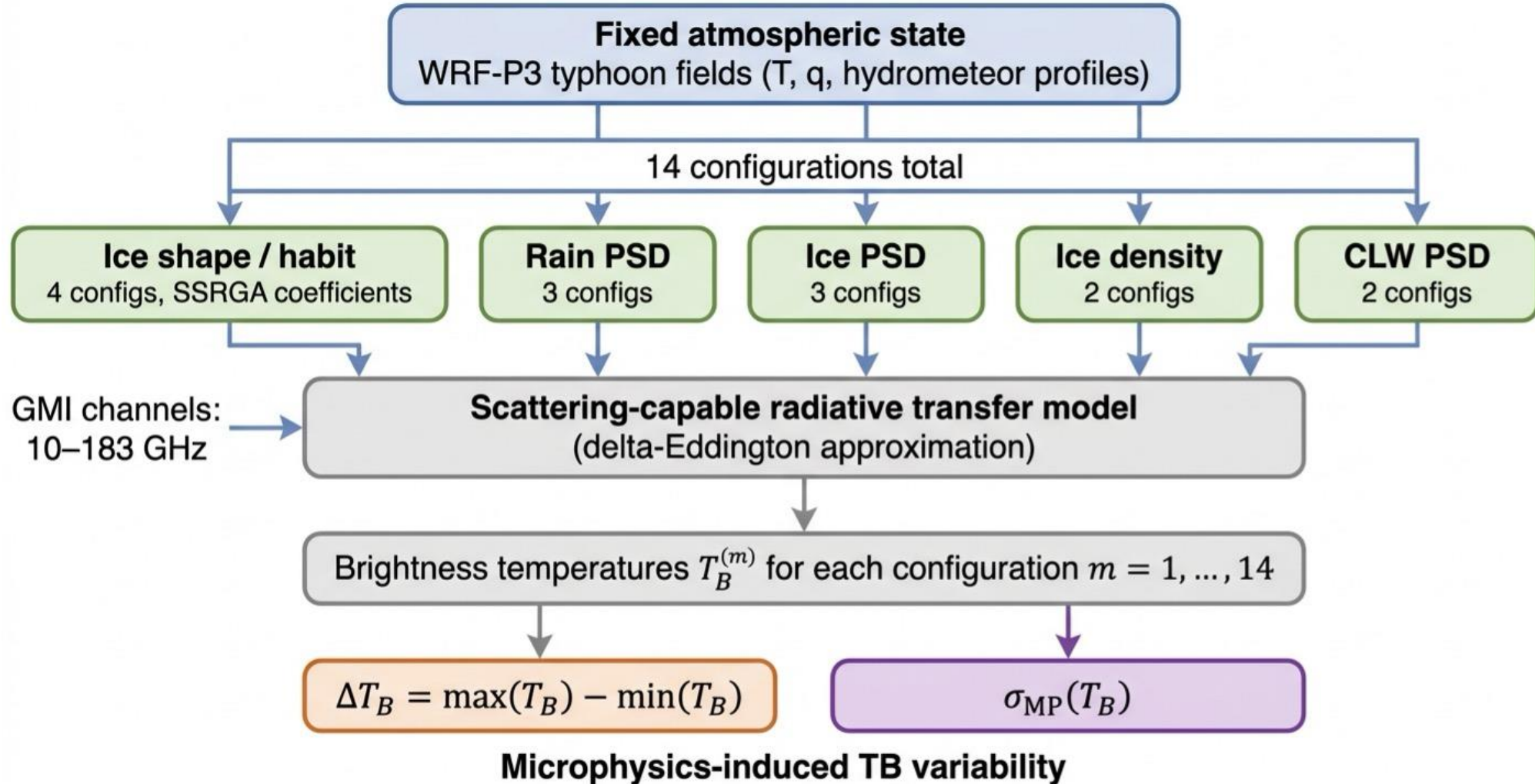
15 typhoon cases — each applied to 14 MP-perturbed sets → 210 total simulations (15 × 14)

CASE#	Scene Name	Location [km]	MSLP [hPa]	MWS [m/s]
CASE#1	SURIGAE_041503	45.95	1.24	6.07
CASE#2	SURIGAE_041615	33.45	-1.13	1.82
CASE#3	SURIGAE_041815	19.48	1.6	-2.42
CASE#4	SURIGAE_042001	25.12	8.43	-7.13
CASE#5	SURIGAE_042300	33.23	0.27	-0.39
CASE#6	NIDA_080506	54.76	3.51	1.50
CASE#7	NIDA_080516	31.30	3.75	1.67
CASE#8	CHANTHU_090921	18.68	-1.91	3.26
CASE#9	CHANTHU_091121	31.55	7.69	-4.08
CASE#10	MINDULLE_092403	55.28	-5.91	5.07
CASE#11	MINDULLE_092902	22.11	11.90	-6.75
CASE#12	MINDULLE_092915	17.40	10.21	-5.91
CASE#13	MALOU_102617	22.43	-4.71	6.29
CASE#14	NYATOH_120207	26.76	-6.66	5.11
CASE#15	RAI_121717	33.38	-0.13	-0.58

Workflow Summary

1 2 3

Fixed-State Radiative-Transfer Experiment Design

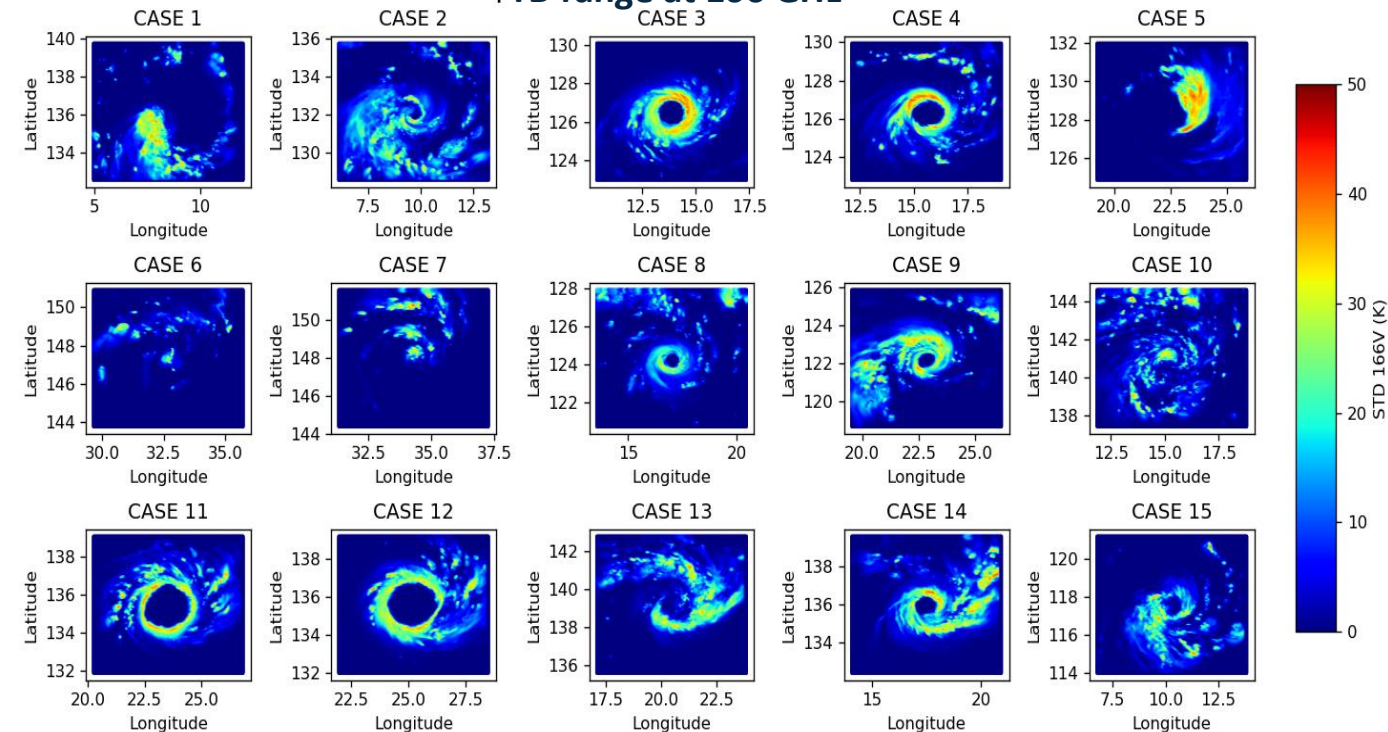


MP assumptions generate structured TB variability

Example: Full TB variability

Standard deviation across 14 MP-perturbed sets, shown for each of the 15 typhoon cases

TB range at 166 GHz



What this shows

Same atmospheric fields, different MP settings
→ spatially organized TB spread.

not random noise

channel dependent

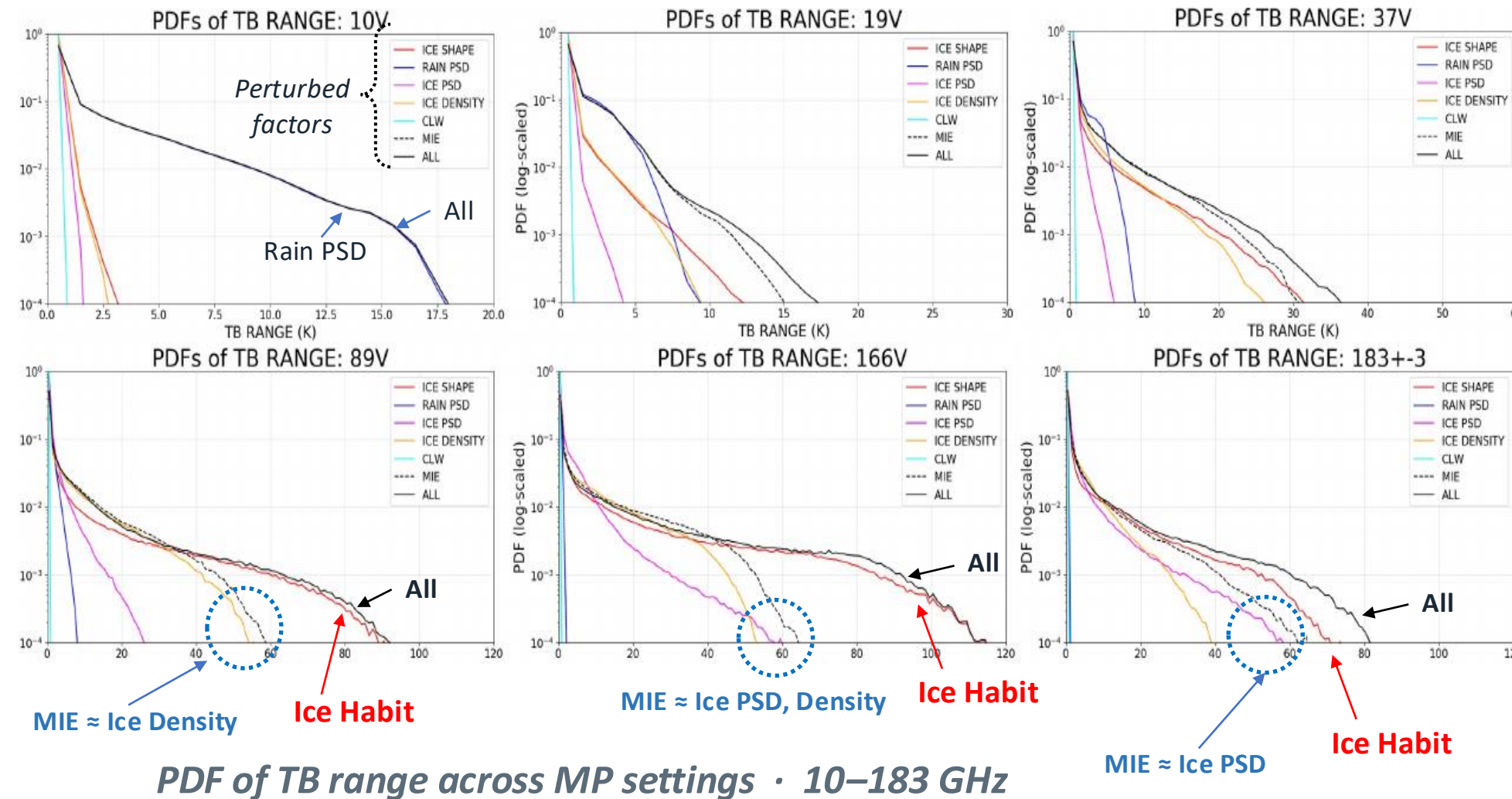
cloud-regime dependent

Bias-correction implication again

A correction model that lacks MP information may absorb this variability into unrelated predictors.

How Much Spread, and From What?

PDFs of TB range by perturbed factors — the dominant factor shifts with frequency



10 GHz

ALL $\approx$ MIE $\approx$ RAIN PSD —
no shape effect; rain PSD
dominates

Higher freq (89-183GHz)

ALL $>$ MIE $\approx$ ICE HABIT —
habit becomes the big
lever

Shape off (Mie case)

89 GHz: MIE $\approx$ ICE
DENSITY dominates
166 GHz: MIE $\approx$ ICE PSD+
DENSITY dominates
183 GHz: MIE $\approx$ ICE PSD
dominates

Predictor Diagnosis

XGBoost learns the MP-driven TB variability from 22 model-based predictors;
SHAP attributes it — organized into five physical families

Hydrometeor amount

*Column-
integrated paths*

IWP
LWP
RWP
CWP

Thermodynamic / moisture

*Temperature &
vapor*

TPW
T2
Z_FREEZE
TPWFRZ

Vertical structure

*Layer heights &
tops*

Z_ICE
Z_RAIN
ZTOP_ICE
ZTOP_RAIN
IWPFRZ

Microphysical property

*Size · density ·
rime state*

D_RAIN_COL
D_ICE_COL
RHO_ICE_COL
RF_ICE_COL

Dynamical

*Surface &
updraft flow*

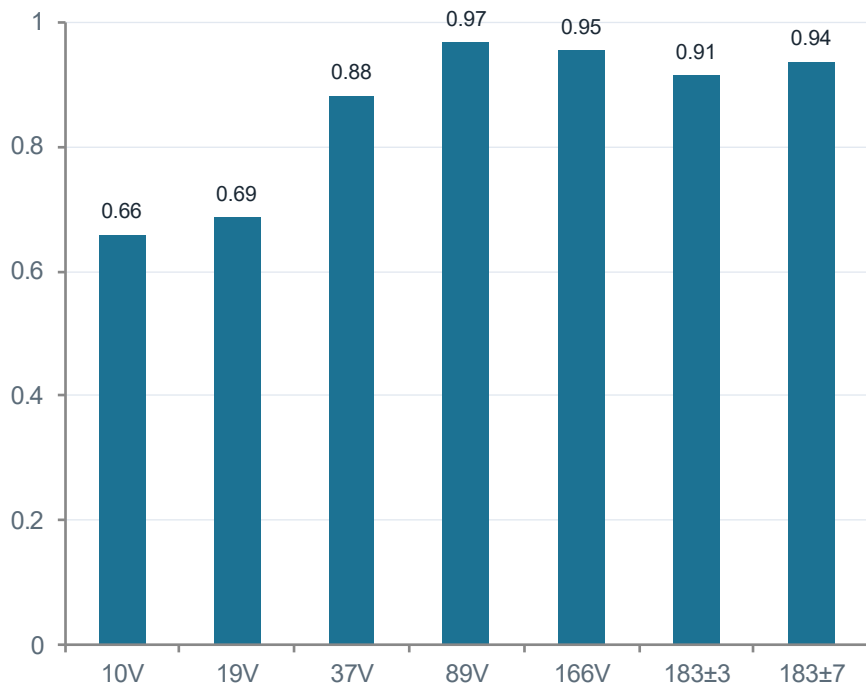
WSPD10
WMAX

22 model-based predictors → XGBoost per channel → SHAP attribution → which family drives TB variability at each frequency

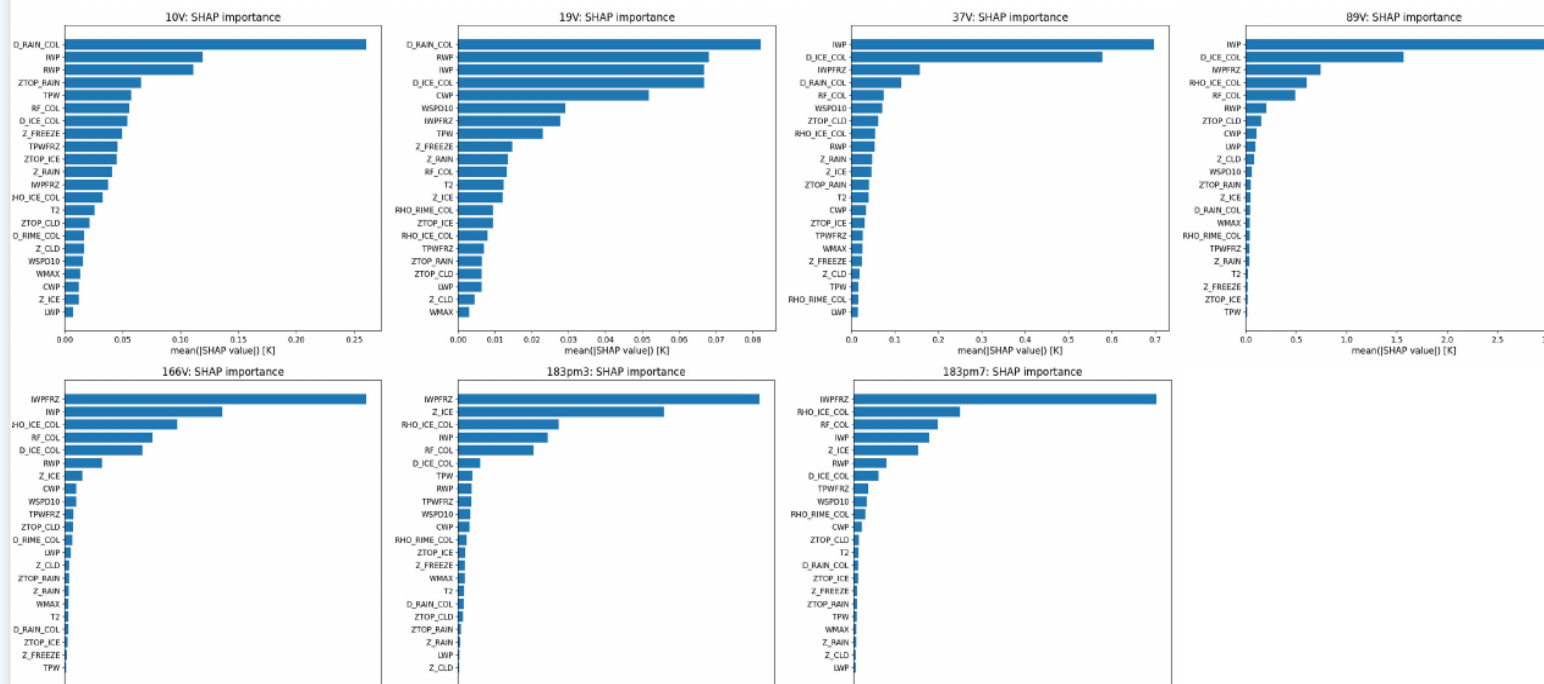
Which Variables Drive the Spread?

XGBoost predicts TB STD from 22 model-based predictors; SHAP attributes each predictor's contribution

XGBoost skill (R^2) — best in scattering channels



Mean |SHAP| importance per channel



10 & 19 GHz

DRAIN_COL rain drop size

37 & 89 GHz

IWP ice water path

166 & 183 GHz

IWPFRZ ice mass above freezing level

The Predictors That Matter

Ranked WRF-output predictors of TB variability — ice mass quantities lead, esp. at scattering channels

 10 & 19 GHz	 37 & 89 GHz	 166 & 183 GHz
1 D_RAIN_COL <i>rain drop size</i>	1 IWP <i>ice water path</i>	1 IWPFRZ <i>IWP above freezing level</i>
2 RWP <i>rain water path</i>	2 D_ICE_COL <i>ice particle size</i>	2 RHO_ICE_COL <i>ice density</i>
3 IWP <i>ice water path</i>	3 IWPFRZ <i>ice mass aloft</i>	3 RF_ICE_COL <i>rime fraction</i>
4 D_ICE_COL <i>ice particle size</i>	4 RHO_ICE_COL <i>ice density</i>	4 IWP <i>ice water path</i>
5 TPW <i>water vapor path</i>	5 RF_ICE_COL <i>rime fraction</i>	5 Z_ICE <i>ice centroid height</i>

Note: RF_ICE_COL, RHO_ICE_COL, D_ICE_COL are available only from WRF-P3 (predicted ice properties) — the physics that six-class schemes cannot expose.

A Missing Term in Bias Correction

1

2

3

Correcting only air-mass and scan bias lets MP variability alias onto them
— We need to add an explicit MP term.

Proposed Microphysics-aware Observation Bias Model

$$b_o = \beta_0 + \underbrace{\sum_i \beta_i^{AM} P_i^{AM}}_{\text{air-mass}} + \underbrace{\sum_j \beta_j^{SCAN} P_j^{SCAN}}_{\text{Scan-geometry}} + \underbrace{\sum_k \beta_k^{MP} P_k^{MP}}_{\text{Microphysics}}$$

$$P^{AM} = \{TPW, T_s, \Gamma, T_{850}, T_{200}\}$$

$$P^{SCAN} = \{\theta, \text{Scan Position}, FOV\}$$

$$P^{MP} = \{IWP, IWP_{FRZ}, D_{ice}, \rho_{ice}, F_{rime}, \rho_{rime}, \dots\}$$

Problem: Microphysical Aliasing

Conventional basis has no **MP bias** term → it aliases onto air-mass and scan biases.



Extension of the Observation Bias Model

- Treat MP assumptions and inconsistency as a structured bias source.
- Use predictors with physical meaning
- Estimate adaptively by channel
 - more investigation for surface, cloud regime.

Thank you

Microphysics-aware RT — from per-channel drivers to smaller, better-targeted O-B biases

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Yonsei University | 12th IPWG Workshop

Questions welcome

APPENDIX

Supporting Detail

Backup slides for Q&A — full setup and per-experiment results.

Forward simulation · Experimental design · EXP#1–#4 · REF vs GMI

Summary: Kim, Shin and Kim (2025), IEEE TGRS: Sensitivity of Microwave Radiative Transfer Simulations to Cloud Microphysics Assumptions.

Methods at a Glance



WRF cloud-resolving runs

12 / 4 / 1.3 km nested, vortex-following; 6 MP schemes (WSM6, WDM6, THOM, MORR, P3-1ICE, P3-2ICE)



5 GPM-GMI typhoon overpasses

Surigae & Mindulle (2021); location error < 120 km at 54 h



Fast plane-parallel RTM

delta-Eddington two-stream; Mie soft-sphere + Maxwell-Garnett; pre-computed bulk-scattering LUTs



Metrics on dTB_x PDFs

$dTB_x = TB_x - TB_{clear}$; PDF-weighted mean & extreme tail; sensitivity = max-min range across assumptions

Five GMI V-pol channels analyzed: 10 · 19 · 37 · 89 · 166 GHz

Experimental Design

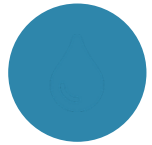
RT SENSITIVITY

1

2

3

Fix the input field (WRF-P3 1ICE), vary one MP assumption at a time in the RTM



EXP#1

Rainwater PSD

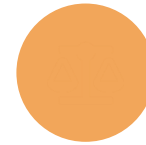
Exponential vs double-moment gamma; vary N_0 , μ



EXP#2

Frozen-particle PSD

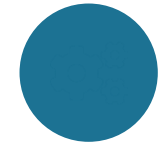
EXPD/DMD + temperature-dependent (Houze, Field)



EXP#3

Frozen density

SPHERE vs COX88 vs MG08 (riming-based)



EXP#4

REF assumptions, fixed field

Six WRF MP schemes in the RTM; isolates the module

REF (reference) = physically consistent: each WRF scheme paired with its own RTM assumptions. **EXP#4 vs REF** isolates the effect of varying input fields.

EXP#1–2 · PSD

$$N(D) = N_0 D^\mu e^{-\lambda D}$$

vary N_0 , μ

EXP#3 · density

$$\rho_{\text{eff}}(D) = \frac{6a}{\pi} D^{b-3}$$

$b < 3 \Rightarrow \rho$ falls with D

Metric — dTB & sensitivity range

$$dT B_x = TB_x - TB_{\text{clear}}$$

$$\text{Range} = \max_m(\overline{dT B_x}) - \min_m(\overline{dT B_x})$$

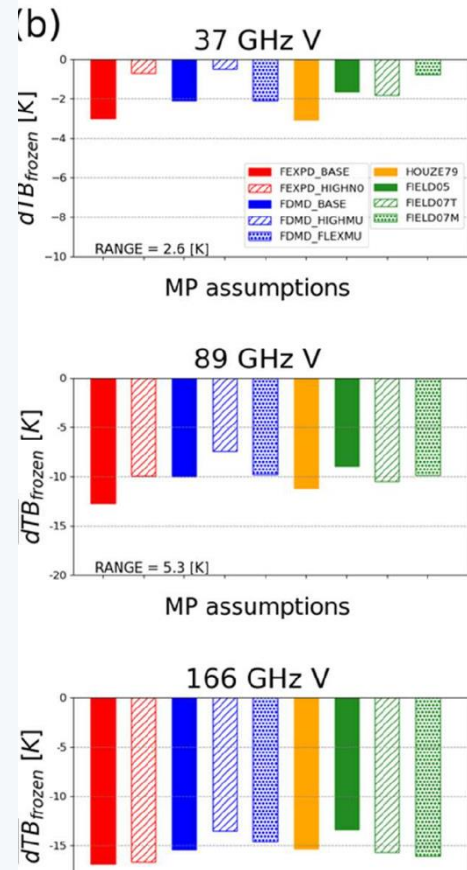
TB depression

vs clear-sky;

Range = spread across MP schemes

EXP#2 Frozen-particle PSD — comparable, then constrained

$\overline{dT}_{B_frozen}$ by MP assumption



PSD families (color = PSD form):

- FEXPD** (*exponential*)
solid / hatched = BASE / HIGHNO
large particles → strongest depression
- FDMD** (*double-moment γ*)
BASE($\mu=0$) / HIGHMU($\mu=8$) / FLEXMU
higher μ → smaller particles → weaker
- HOUZE79** (*empirical*)
N(D) compensates low/high-T
moderate depression
- FIELD** (*empirical (Field)*)
05 / 07T / 07M — super-exponential
strong shift to small → weakest depression



Range 2.6 – 5.3 K

Comparable in magnitude to rainwater PSD (EXP#1)



Small particles win

Super-exponential Field PSDs shift mass to small sizes → weak depression



Forward-scatter cancels

At 166 GHz high g (→0.78) offsets extreme depression



Net effect: constrained

Variability shrinks at high frequency, like rainwater

Paper Fig. 8(b): dTB_{frozen} with RANGE labels (9 MP assumptions)

REF vs GMI Observations

Even the most realistic configuration can give the worst fit

Low frequency (10–37 GHz)

REF under-estimates TB enhancement → points to too-weak WRF rainwater fields.

High frequency (89–166 GHz)

REF under-estimates TB depression → Mie soft-sphere over-estimates g and over-modifies Q_{sca} for non-spherical ice.



The surprise. P3-1ICE / P3-2ICE — with the **most sophisticated density scheme (MG08)** — give the worst high-frequency JSD. A partially realistic component, fed through an unrealistic Mie soft-sphere model, degrades the result more than a simpler but consistent one.

Best Jensen–Shannon match to OBS (case-dependent):

10 GHz

MORR

19 GHz

P3-2ICE

37 GHz

P3-2ICE

89 GHz

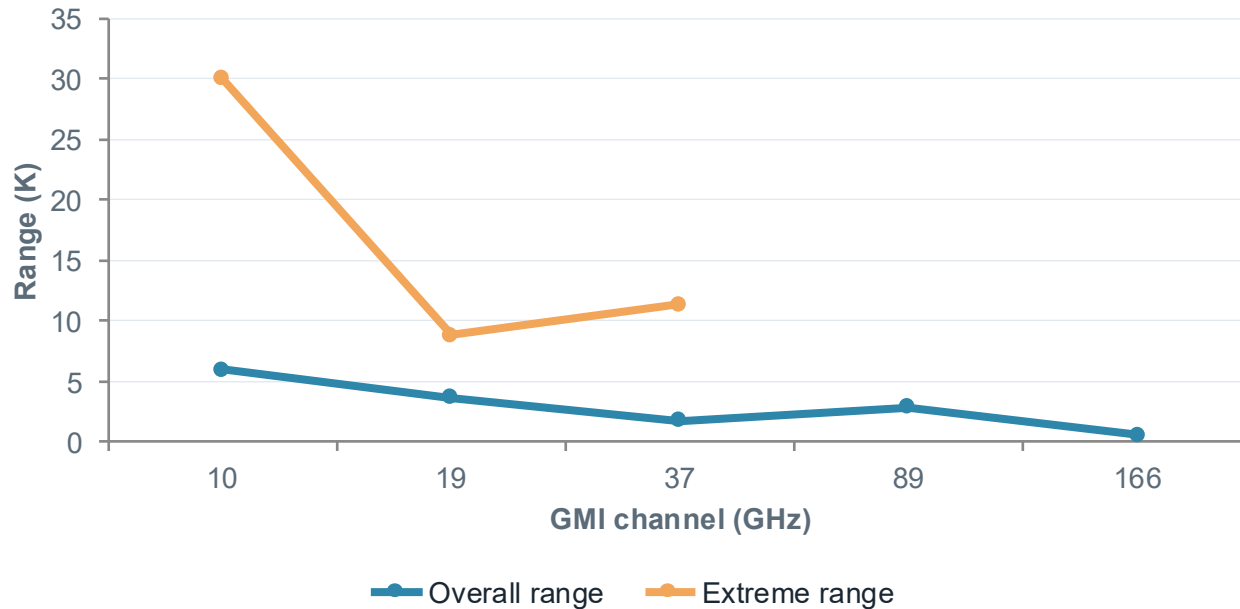
MORR

166 GHz

THOM

PSD: The Self-Limiting Lever

Sensitivity to rainwater PSD assumptions (max-min range, K)



Largest at 10 GHz

6 K overall — large-particle emission dominates rainwater extinction



Collapses with frequency

Down to 0.6 K at 166 GHz



Two offsets at work

Small- vs large-particle extinction + emission vs scattering cancel



EXPD vs DMD $\approx$ minor

Changing N_0 / μ matters more than the distribution family

Density:

SUPPORTING EVIDENCE

1

2

3

Frozen-particle density assumptions produce the largest, most robust TB variations

9.3 K

overall range at 166 GHz

≈3× the PSD effect

49 K

extreme range at 89 GHz

≈250% of the PSD effect

Q_sca

the mechanism

density reshapes scattering efficiency



COX88 (non-spherical, $b < 3$)

↓ Q_sca for large particles → weaker depression at 37/89 GHz;
↑ Q_sca for small particles boosts 166 GHz when paired with flexible DMD.



MG08 (riming-based, P3)

Most realistic density; but P3 fields rarely rime ($F_{\text{rime}} < 0.5$) → weakest scattering of all — a realism-vs-consistency warning revisited later.

Key points: of all MP choices in the scattering module, **density is the one to get right** — it controls Q_sca more than the self-limiting PSD levers. **Realism ≠ accuracy:** the most "realistic" setup can give the worst fit to GMI (REF vs obs).

EXP#4 vs REF — Fields interact with the module

RT SENSITIVITY

1

2

3

Varying the input atmospheric field adds variability and can reinforce or suppress signals

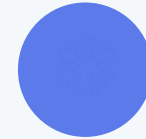


Case A — WDM6 rainwater

WRF-WDM6 makes abundant mid-level rainwater → expected strong 10–19 GHz enhancement.

Suppressed by the RTM's DMD $\mu=1$ assumption, which shifts $N(D)$ to small particles.

$\overline{dTB}$ drops 9.3 → 7 K at 10 GHz.



Case B — THOM frozen

WRF-THOM makes abundant snow; FIELD05 PSD adds small particles; COX88 boosts their Q_{sca} .

Reinforced — the trio compounds into a strong 166 GHz depression.

$\overline{dTB}$ spread 7.7 → 16.4 K (REF), more than double EXP#4.

Lesson: MP assumptions in the NWP field and in the RTM module are not independent — **physical consistency across the whole chain matters.**

Key Takeaways



PSD impact is self-limiting

Rainwater & frozen PSD effects shrink at high frequency via extinction compensation and forward-scatter cancellation.



Density is the dominant lever

Frozen-particle density drives the largest TB variations (up to 9.3 K overall, 49 K extreme) through Q_{sca} .



Consistency over isolated realism

Keep PSD & density assumptions consistent across field, RTM, and scattering DB (DDA databases) — relevant to all-sky DA & cloud-property retrievals.

Microphysics systematically shapes TB

→ but does our DA bias correction even account for it?

Does DA bias correction account for microphysics?



TB VARIABILITY DRIVERS

The predictor set was designed for air-mass and scan errors — not for microphysics.

CONVENTIONAL OBSERVATION-BIAS MODEL (VarBC)

$$b_o = \underbrace{\beta_0}_{\text{offset}} + \underbrace{\sum_i \beta_i^{AM} P_i^{AM}}_{\text{air-mass}} + \underbrace{\sum_j \beta_j^{SCAN} P_j^{SCAN}}_{\text{Scan-geometry}}$$

air-mass predictors: $P^{AM} = \{TPW, T_s, \Gamma, T_{850}, T_{200}\}$

scan-position predictors: $P^{SCAN} = \{\theta, \text{Scan Position}, FOV\}$

Microphysics-induced TB variability — rain & ice PSD, habit, density — has no predictor of its own.
Captured? No — yet not gone: it is aliased into the air-mass & scan coefficients, **quietly contaminating the correction.**

Adding Microphysics (MP) correction: $+\sum_k \beta_k^{MP} P_k^{MP}$? $\rightarrow$ then the predictors?

01. Calculating TB variability arising from various MP assumptions

Various MP settings

*Default: P3 settings

*MGD: Modified-Gamma dist.

	MP name	CLW	RAIN	ICE PSD	ICE Density	ICE shape
1	Default	Single-moment MGD (Martin <i>et al.</i> , 1994)	Double-moment MGD ($\mu = 0$)	Double-moment MGD ($\mu = f(\lambda)$)	Morrison and Grabowski (2008)	MIE
2	ARTS-OPT	Default	Default	Default	Default	ARTS-OPT (Kim <i>et al.</i> 2024)
3	SSRGA-OPT	Default	Default	Default	Default	SSRGA-OPT (See poster 9.15)
4	SSRGA-RIME	Default	Default	Default	Default	SSRGA-RIME (See poster 9.15)
5	RAIN_WSM6	Default	Single-moment MGD	Default	Default	Default
6	RAIN_WDM6	Default	Double-moment MGD ($\mu = 1$)	Default	Default	Default
7	ICE_PSDWSM6	Default	Default	Houze <i>et al.</i> (1979)	Default	Default
8	ICE_PSDTHOM	Default	Default	Field <i>et al.</i> (2005)	Default	Default
9	ICE_PSDMORR	Default	Default	Double-moment MGD ($\mu = 0$)	Default	Default
10	ICE_DENWSM6	Default	Default	Default	Spherical	Default
11	ICE_DENTHOM	Default	Default	Default	Cox (1988)	Default

01. Calculating TB variability arising from various MP assumptions

Various MP settings

*Default: P3 settings

*MGD: Modified-Gamma dist.

	MP name	CLW	RAIN	ICE PSD	ICE Density	ICE shape
#1	Default	Single-moment MGD (Martin <i>et al.</i> , 1994)	Double-moment MGD ($\mu = 0$)	Double-moment MGD ($\mu = f(\lambda)$)	Morrison and Grabowski (2008)	MIE
12	CLW_WSM6	Monodisperse	Default	Default	Default	Default
13	CLW_THOM	Thompson <i>et al.</i> (2008)	Default	Default	Default	Default
14	CLW_MORR	Curry (1999)	Default	Default	Default	Default

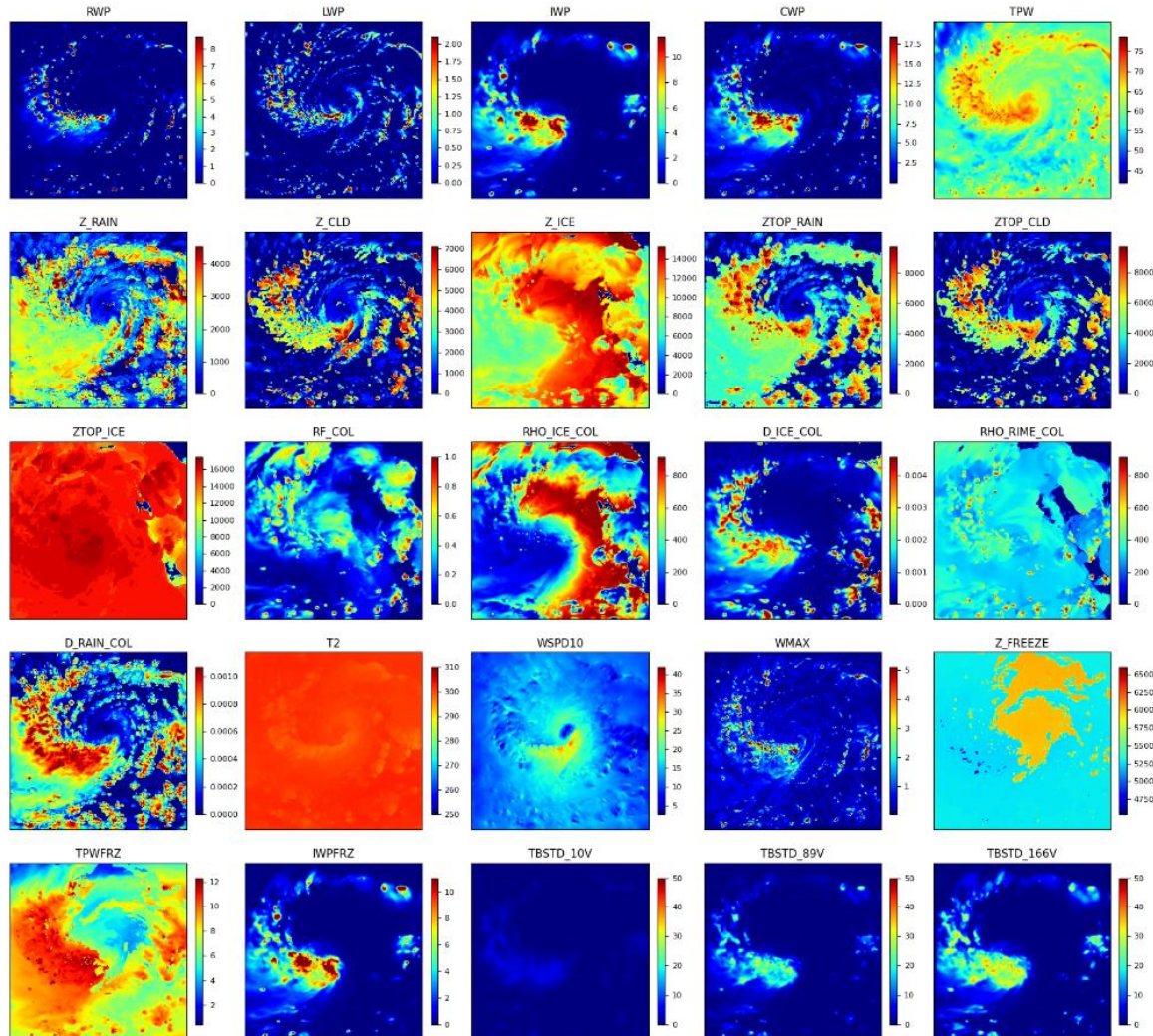
- Set#1: [Default, #2-4] => Shape effect
- Set#2: [Default, #5-6] => RAIN PSD effect
- Set#3: [Default, #7-9] => ICE PSD effect
- Set#4: [Default, #10-11] => ICE Density effect
- Set#5: [Default, #12-14] => CLW PSD effect
- MIE: [Default, #5-14] => Exclude Shape-related variability (assume MIE)
- ALL => represent full TB variability

Predictors of TB variability

Model-based predictors

22 Possible predictors of TB variability (*i.e.*, TBSTD)

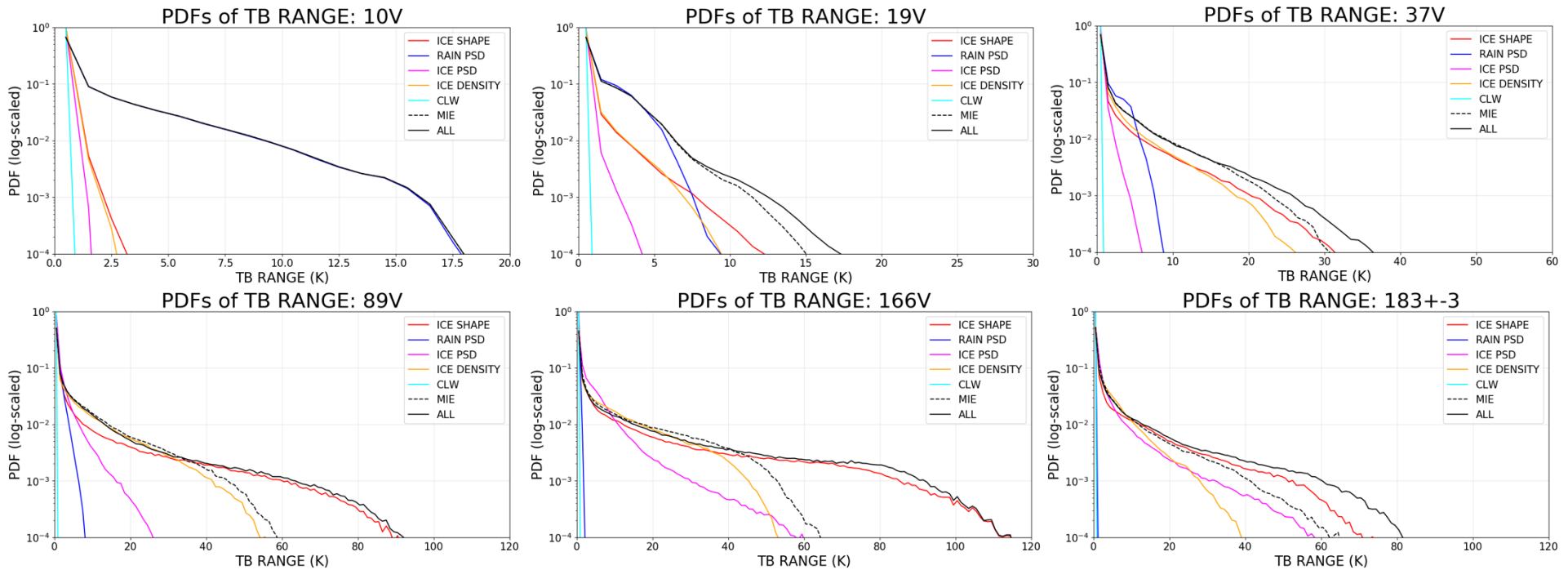
CASE 1: Predictors and TB STD



- IWP: ICE Water Path
- LWP: CLW Path
- RWP: RAIN Water Path
- CWP: IWP+LWP+RWP
- TPW: Water Vapor Path
- Z_X: mass-weighted average altitude of hydrometeor X
- ZTOP_X: top altitude of hydrometeor X
- RF_ICE_COL: mass-weighted average F_{rime}
- RHO_ICE_COL: mass-weighted average ρ_{ice}
- D_ICE_COL: mass-weighted average D_{ice}
- RHO_ICE_COL: mass-weighted average ρ_{rime}
- $*F_{rime}, \rho_{ice}, D_{ice}, \rho_{rime}$ are only available from WRF-P3 simulations
- D_RAIN_COL: mass-weighted average proxy of rain diameter
- T2: 2-m Temperature
- WSPD10: 10-m Wind Speed
- WMAX: maximum vertical velocity
- Z_FREEZE: Freezing level height
- TPWFRZ: TPW above Z_FREEZE
- IWPFRZ: IWP above Z_FREEZE

01. Calculating TB variability arising from various MP assumptions

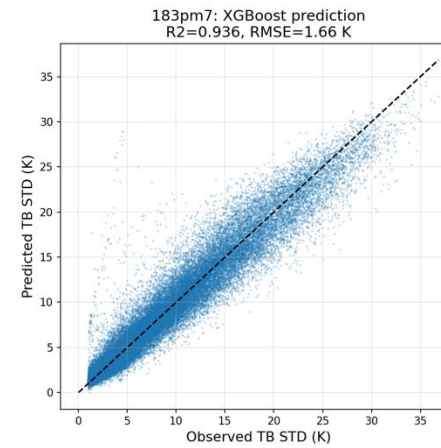
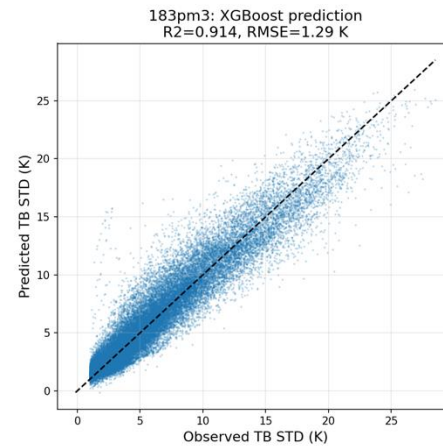
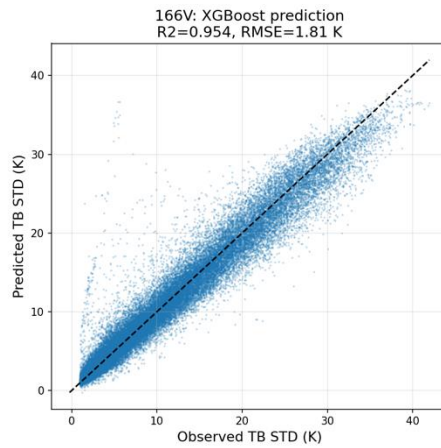
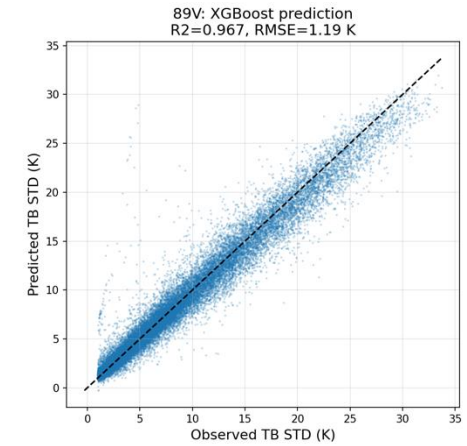
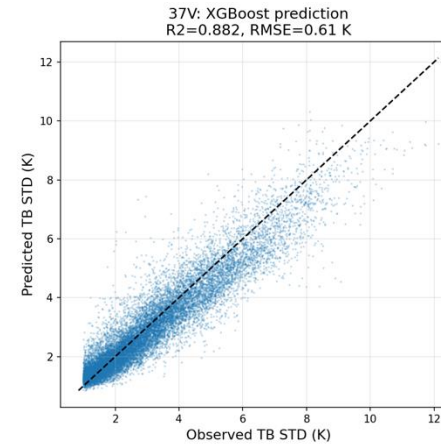
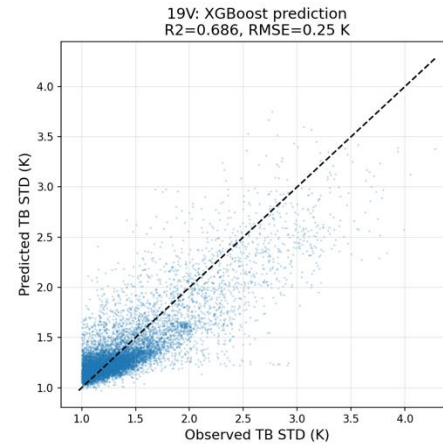
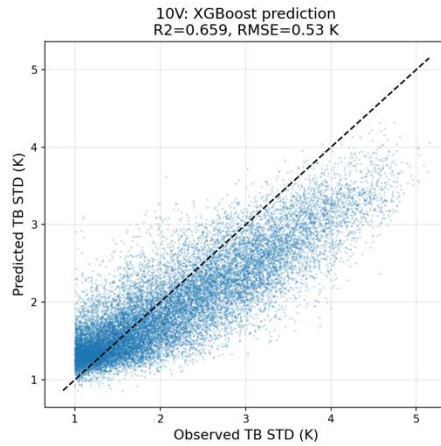
Distribution of TB range



- 10 GHz: TB range in ALL = MIE ~ RAIN PSD => No shape effect, large RAIN PSD effect
- Higher frequency: TB range in ALL > MIE ~ ICE SHAPE => large shape effect
- 89 GHz: TB range in MIE ~ ICE DENSITY => when shape effect excluded, large Density effect
- 89 => 183 GHz: TB range in MIE ~ ICE PSD => when shape effect excluded, large PSD effect

03. Finding drivers of TB variability

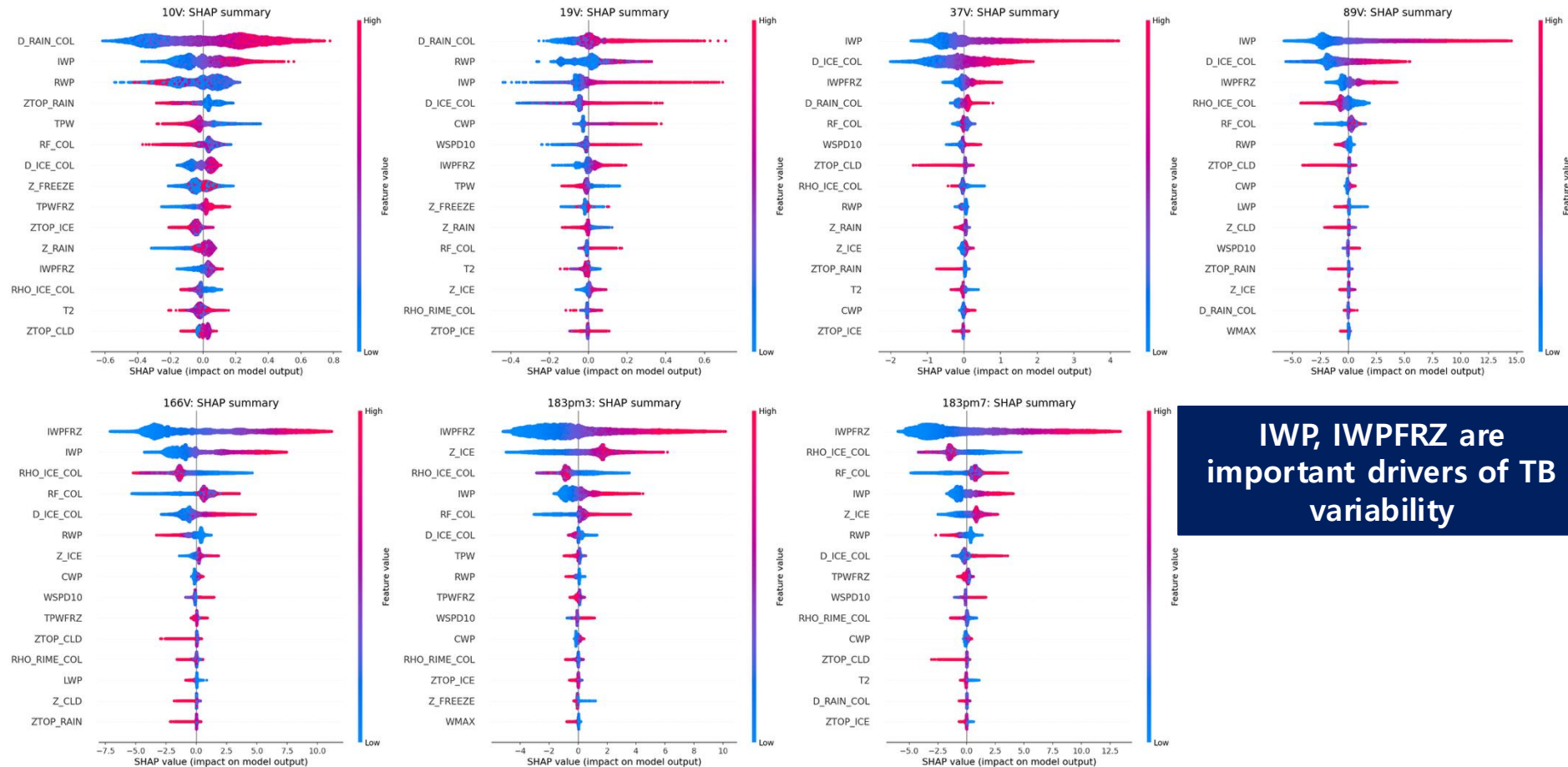
XGBoost Prediction



**Better performance at
scattering channels
(Best at 89 GHz V)**

03. Finding drivers of TB variability

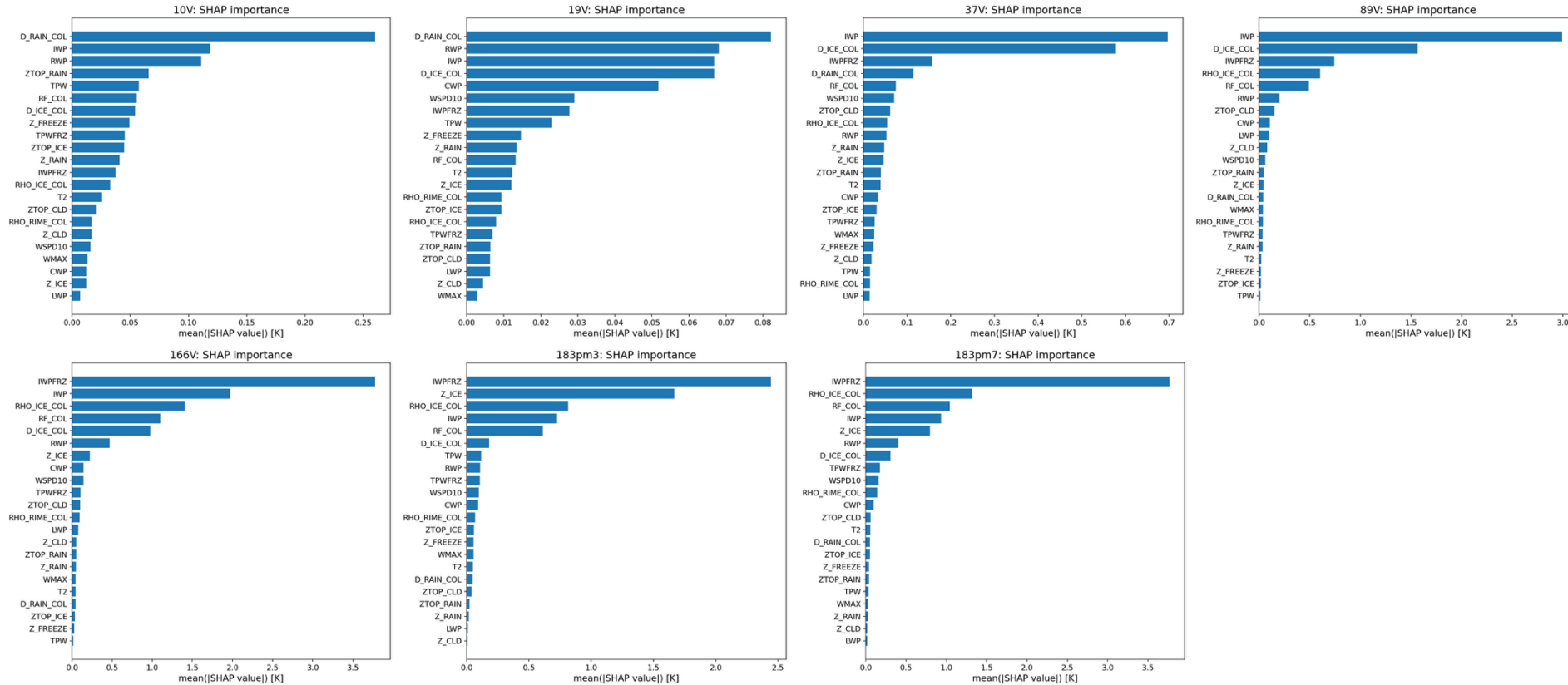
SHAP values



IWP, IWPFRZ are important drivers of TB variability

03. Finding drivers of TB variability

Mean SHAP values



DRAIN_COL (10 and 19 GHz), IWP (37 and 89 GHz), IWPFRRZ (166 and 183 GHz), are primary drivers of TB variability